

3.2 AN OVERVIEW OF MAP/GLOBUS NO<sub>x</sub>

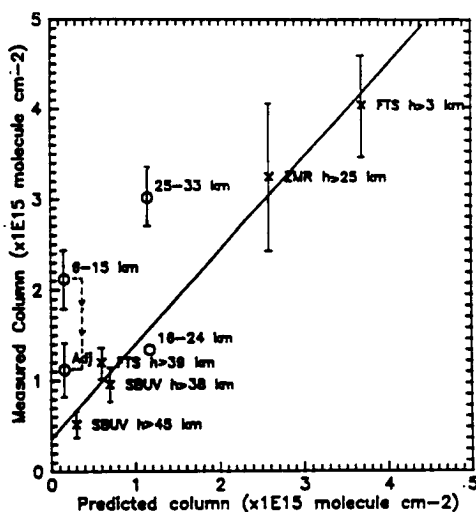
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GLOBUS NO<sub>x</sub> is a combined observation of nitrogen compounds by more than 20 experiments, from satellites, remote and in situ instruments on board balloons and ground observatories, within a short time period. Held in September 1985 above Southern France, the field campaign has been a technical success. Most of the observations have been achieved as anticipated. NO, NO<sub>2</sub> and relevant species and physical parameters involved in their photochemistry, were measured between 5 km and 40 km at several periods of the day. A first step of data interpretation which consists of instrumental intercomparisons, is now achieved. Several systematic biases between data of various origins which have appeared in the past, are now understood and reduced. A second step which deals with atmospheric photochemistry issues like diurnal cycles and budget, is now on its way. It will be the object of a close exchange between experimenters involved in the campaign and modelers.

TABLE 1. Measurements Achieved During GLOBUS NO<sub>x</sub>

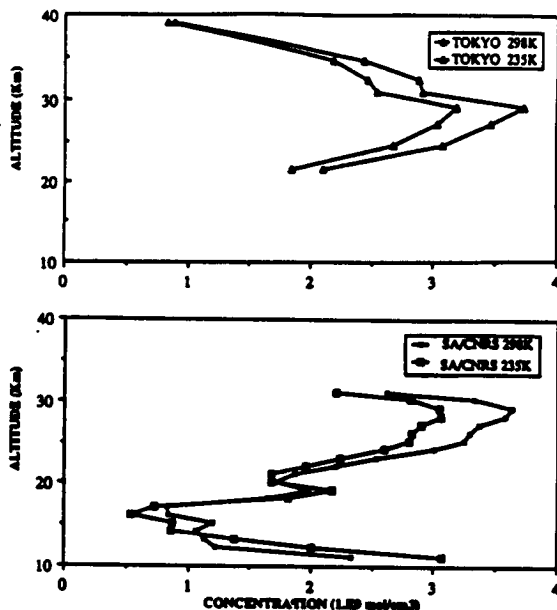
	Balloon	Satellite	Ground	Total
NO <sub>x</sub>	11	3	6	20
Other	11	3	5	19
Total	22	6	11	39



### Comparison of NO Measurements

- Ground-based ZMR, FT spectrometer
- Balloon-borne IR interferometer
- SBUV
- In situ chemiluminescence.

Figure 1. Intercomparison of NO column measurements during MAP/GLOBUS 1985, R. L. McKenzie, W. A. Matthews, Y. Kondo, R. Zander, P. Demoulin, P. Fabian, D. G. Murcray, F. J. Murcray, O. Lado-Bordowsky, C. Camy-Peyret, H. K. Roscoe, J. A. Pyle, R. D. McPeters, submitted to *J. Atmos. Chem.*, 1988.



### NO<sub>2</sub> Remote Sensing Comparison: Solar Visible Occultation from Balloon

- Three instruments on a common balloon gondola (Tokyo, IASB, CNRS)
- Method overestimates NO<sub>2</sub> concentration because of temperature dependence of the absorption cross sections of the gas. Correcting factor between 12% and 18% in the stratosphere
- Total uncertainty: 15% in the stratosphere, 30% in the troposphere

Figure 2. Balloon observations of nitrogen dioxide by visible occultation during GLOBUS NO<sub>x</sub>, J. P. Pommereau, F. Goutail, N. Iwagami, K. Shibasaki, P. C. Simon, W. Peetermans, J. P. Naudet, P. Rigaud, D. Huguenin, submitted to *J. Atmos. Chem.*, 1988.

**NO<sub>2</sub> Remote Sensing Comparison:  
SAGE 2 and Balloon Occultation**  
 -Differences much smaller than Instruments  
 Uncertainties  
 -Average systematic difference < 4%

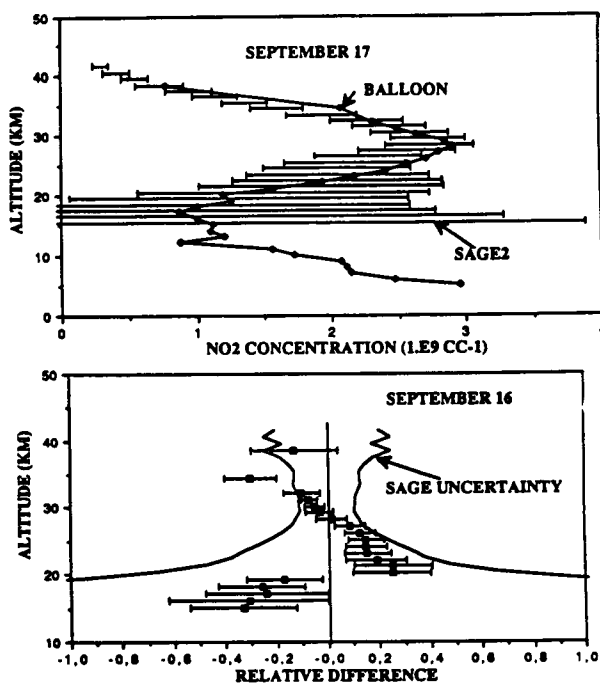
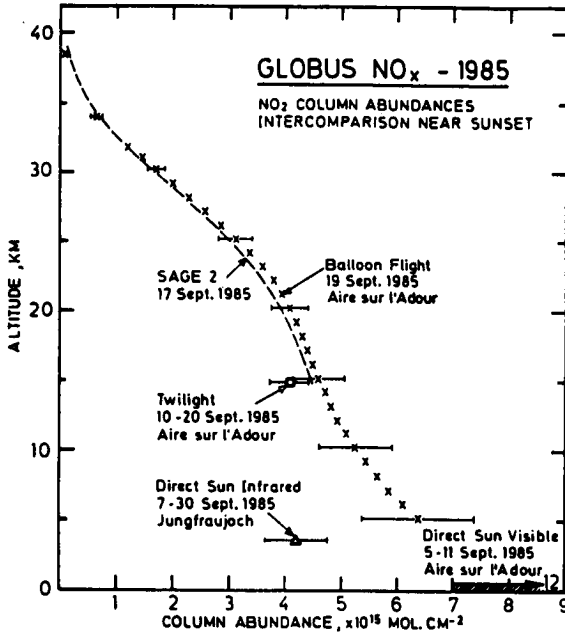


Figure 3. Nitrogen dioxide data comparison with infrared and visible balloon remote sensing measurements during GLOBUS NO<sub>x</sub>, J. Lenoble, W. P. Chu, D. G. Murcray, O. Lado-Bordowski, C. Camy-Perret, A. Perrin, J. P. Pommereau, F. Goutail, K. Shibasaki, N. Iwagami, and P. C. Simon, submitted to Quad. Ozone Symp., 1988.



### Comparison of Ground-Based, Balloon and Satellite NO<sub>2</sub> Measurements

- Data generally consistent when local NO<sub>2</sub> contamination inside the boundary layer absent
- Further simultaneous direct sun IR and visible observations needed

Figure 4. Measurements of column abundances of nitrogen dioxide, NO<sub>2</sub>, from the ground during the GLOBUS NO<sub>x</sub> campaign, R. Zander, P. Demoulin, G. Roland, W. A. Matthews, P. V. Johnston, J. P. Pommereau, N. Iwagami, K. Shibasaki, submitted to Quad Ozone Symp., 1988.

### Ozone Measurements 30-55 km

- Microwave at Bordeaux Observatory
- Umkehr at Haute-Provence Observatory
- Satellite EXOS-C BUVa

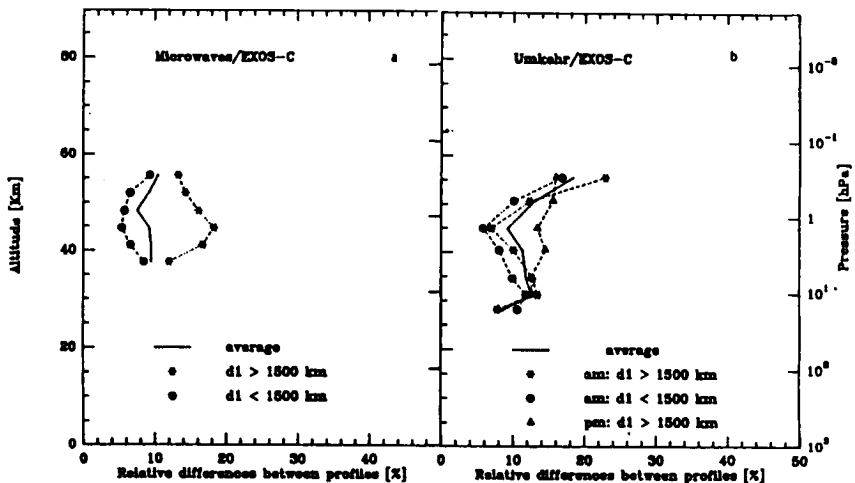


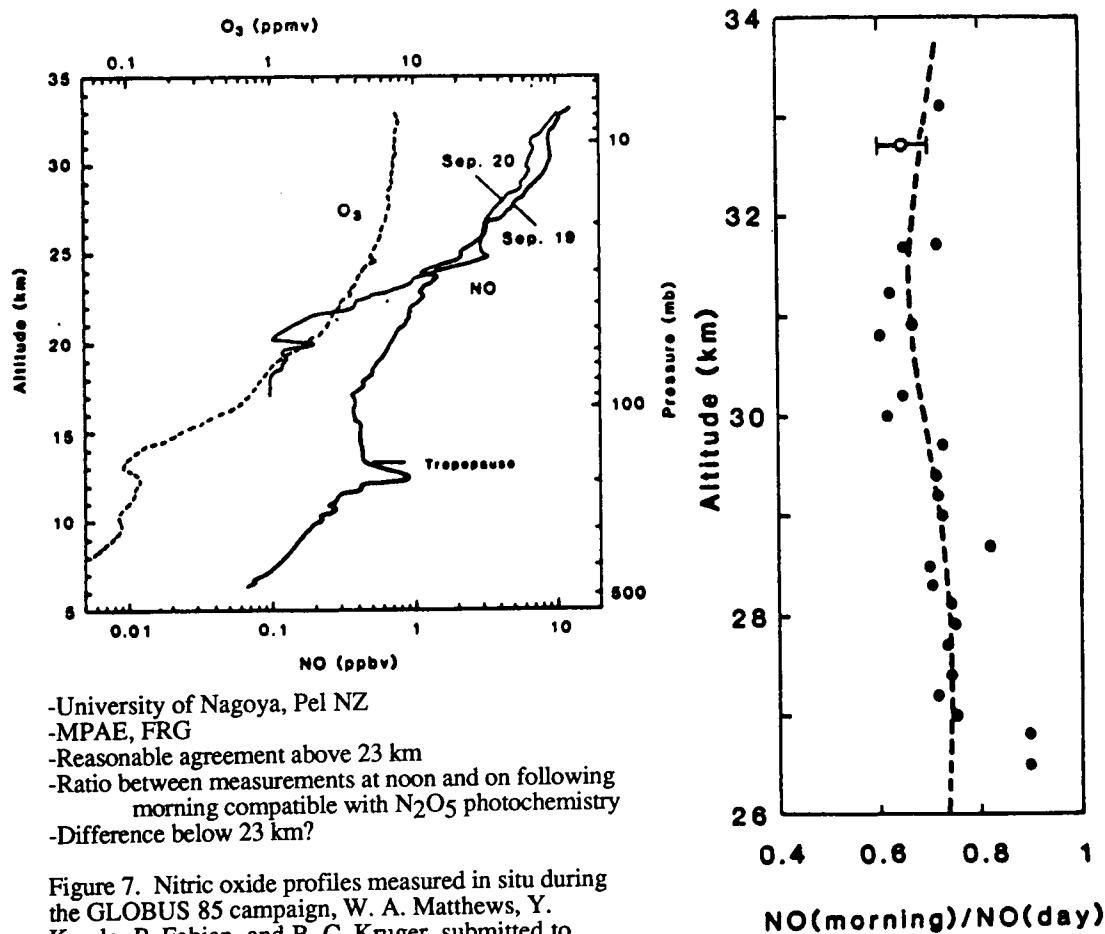
Figure 5. Comparison of stratospheric and mesospheric ozone profiles obtained by ground-based and satellite observations, J. de La Noe, M. Koike, T. Ogawa, P. Marché, submitted to Quad Ozone Symp., 1988.

### Ozone Instruments Total Accuracy

- In situ UV Photometer: 3% Total Uncertainty
- Remote Visible Occultation: 5-10%
- SAGE II: 10%
- In situ Gas Phase Chemiluminescence: 10%
- Umkehr: Layers 4 and 5, 10%; Layer 6, 15%
- SBUV: 10% Average, Peak Systematically Lower
- EXOS-C BUV: 10-15%
- Brewer-Mast Sondes: 5-10% between 18 and 25 km, 15-20% above
- Solar UV Occultation: Systematic Underestimation by 30-35%

Figure 6. Measurements of stratospheric ozone during the MAP/GLOBUS NO<sub>x</sub> campaign, D. Robbins, P. Amedieu, J. Pelon, J. P. Pommereau, F. Goutail, N. Iwagami, K. Shibasaki, T. Ogawa, M. Koike, P. Marché, J. P. Naudet, P. Rigaud, D. Huguenin, J. Lenoble, G. Maddrea Jr., submitted to *J. Geophys. Res.*, 1988.

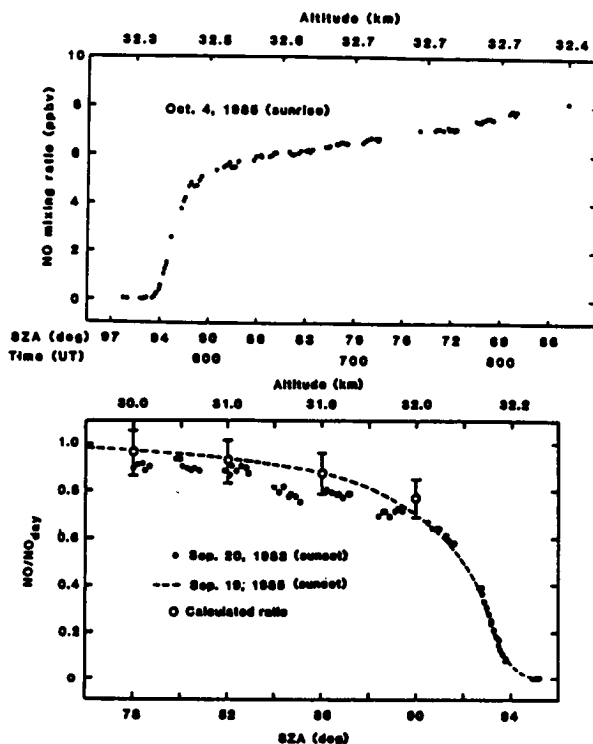
### NO Vertical Profiles by In situ Chemiluminescence



- University of Nagoya, Pel NZ
- MPAE, FRG
- Reasonable agreement above 23 km
- Ratio between measurements at noon and on following morning compatible with N<sub>2</sub>O<sub>5</sub> photochemistry
- Difference below 23 km?

Figure 7. Nitric oxide profiles measured in situ during the GLOBUS 85 campaign, W. A. Matthews, Y. Kondo, P. Fabian, and B. C. Kruger, submitted to *J. Atmos. Chem.*, 1988.

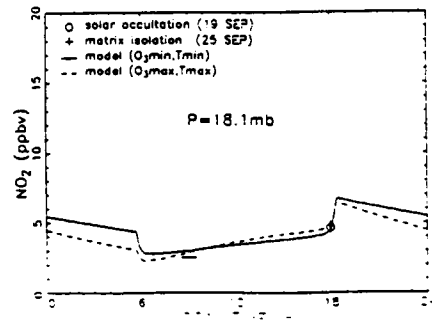
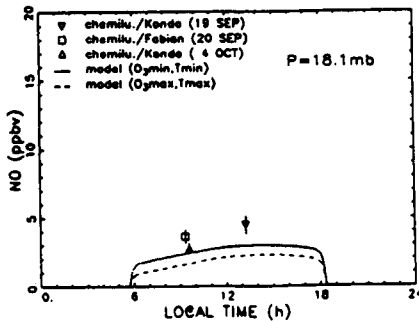
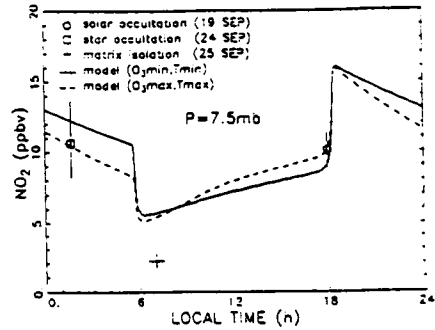
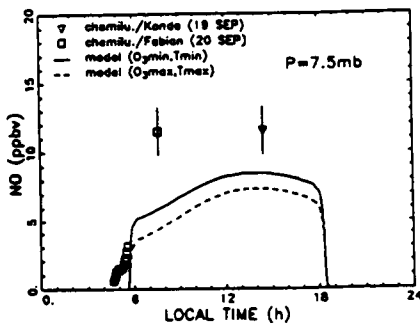
# NO Diurnal Variation



- Variation observed in situ from sunrise until sunset around 32 km with two balloon flights
- From rate of increase of NO during morning:  $[N_2O_5]$  at the end of the night =  $3.9 \pm 1.5$  ppbv
- NO decrease at the end of afternoon at  $SZA > 80^\circ$ , requires decrease of  $JNO_2$

Figure 8. Diurnal variation of nitric oxide at 32 km: measurements and interpretation, Y. Kondo, W. A. Matthews, P. Amedieu, D. E. Robbins, *J. Geophys. Res.*, in press, 1988.

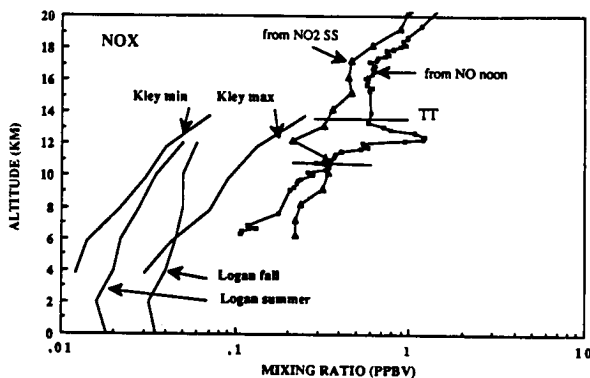
# NO/NO<sub>2</sub> Diurnal Variation. Model Comparisons



- Agreement between NO<sub>2</sub> Predicted and Measured (Night, Morning, Evening)
- Much More NO than Predicted, Partitioning between NO and NO<sub>2</sub> differs from that Observed, Difference Increases at High Altitude
- NO decay in the Afternoon at SZA > 80°. JNO<sub>2</sub> Reduced by Multiple Scattering?

Figure 9. Diurnal variation of stratospheric NO and NO<sub>2</sub> from MAP/GLOBUS 1985, J. P. Naudet, M. Pirre, R. Ramaroson, P. Rigaud, P. Fabian, M. Helten, N. Iwagami, K. Shibasaki, T. Ogawa, Y. Kondo, W. A. Matthews, J. P. Pommereau, F. Goutain, P. C. Simon, W. Peetermans, submitted to Quad. Ozone Symp., 1988.

### A $\text{NO}_x$ Source in the Upper Troposphere?



- $\text{NO}_x$  Observed Larger than Predicted by Factor 3 to 10
- Models Include Mean Transport from Stratosphere
- Local Transport from 20 km at Least, Would Have Left Signature on Potential Temperature and Ozone
- Transport from Surface Level Would Require Much More than  $\text{NO}_x$  Lifetime [Kley et al., 1981]

Conclusion: Source present in Upper Troposphere

Figure 10. Observed and predicted  $\text{NO}_x$  below 20 km [Logan et al., 1981; Kley et al., 1981]. A  $\text{NO}_x$  source in the upper troposphere?, J. P. Pommereau, F. Goutail, Y. Kondo, W. A. Matthews, M. Helten, submitted to Quad. Ozone Symp., 1988.

## CONCLUSIONS

### Experimental

- Better evaluation of  $\text{NO}$ ,  $\text{NO}_2$ ,  $\text{O}_3$  experimental uncertainties
- Instruments improvements: Result of previous campaigns
- Systematic errors found and reduced
- Consistency of all  $\text{NO}$  data after correction of chemiluminescence data?
- Further comparison between IR and visible ground-based  $\text{NO}_2$  instruments still required
- Interpretation of  $\text{NO}_x$  measurements by chemiluminescence and converter still to come

### $\text{NO}_x$ Photochemistry

- Excess of  $\text{NO}$  compared to  $\text{NO}_2$  in the stratosphere. Significant? Mechanism?
- $\text{NO}$  decrease in the afternoon after SZA  $80^\circ$ .  $J_{\text{NO}_2}$  decrease? Effect of multiple scattering?
- Source of  $\text{NO}_x$  in the upper troposphere. Aircraft emission? Lightning?
- Interpretation will follow on. Data base ( $\text{NO}_x$  constituents, related species and atmospheric parameters) available for further studies